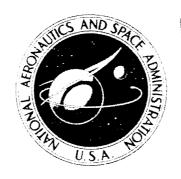
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DESIGN STUDIES ON THE EFFECTS
OF ORIENTATION, LUNATION, AND LOCATION
ON THE PERFORMANCE OF LUNAR RADIATORS

by Thomas Dallas, Anthony J. Diaguila. and James F. Saltsman

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SUMMARY

The performance of lunar radiators was studied for various radiator orientations (with respect to the lunar surface and the Sun), lunar day and night conditions, and location.

The vertical radiator with two active sides had a higher heat-rejection rate than the horizontal radiator of the same overall dimensions with one active side. Direct solar radiation has little effect on the performance if the solar absorptance can be made small (about 0.20 or less) for low radiator temperatures (about 760° R or 422 K). For higher radiator temperatures (1460° R or 811 K), solar radiation effects remain small, even with solar absorptance values of 0.75.

For low-temperature cooling requirements, a horizontal radiative surface is essential. A radiator in polar location would be appreciably smaller because the Moon surface is colder and there is no sunshine.

Graphical solutions of the design equations are presented. This expedites the procedure for approximating prime radiator areas for preliminary studies of lunar heat-rejection systems. The effective sink temperature for any given set of radiator conditions is readily obtained from equations and graphs are presented.

INTRODUCTION

In space, heat can be rejected from a system only by radiant heat transfer. For this reason, a space radiator might be quite large unless high radiating temperatures can be used. Large areas plus additional material required for meteoroid protection result in heavy waste-heat-rejection systems.

The sizes of space radiators, whether they are for heat-rejection systems of a space or lunar power plant, for cooling of the power plant components, or for refrigeration sys-

tems, are basically determined by the amount of heat they must reject, temperature of the radiator surface, and the effective sink temperature. The radiator temperature is essentially determined by thermodynamic cycle efficiency, pumping fluid characteristics, and allowable material temperatures. In lunar radiators, the effective sink temperature is influenced by the orientation of the radiator surface with respect to the Sun and the surface of the Moon, as well as daytime and nighttime conditions and radiator location. The performance of a radiator is improved by lower sink temperatures which results in smaller radiator areas.

Design analyses for space and lunar power systems have shown that the waste-heat radiators for these systems are, in many cases, the heaviest components in the complete power system. This study analyzed the effects of radiator orientation, location, and lunation on the heat-rejection capabilities of lunar radiators.

Heat-rejection rates of a radiator operating over a complete lunation (which is the interval from one new moon to the next) are presented for three specific cases:

Case A - Radiating surface flat on the equator of the Moon with solar radiation incident on its surface. The side facing the Moon is insulated.

Case B - Radiating surface upright on the equator of the Moon with solar radiation incident on its surface. The plane of the radiator is normal to the plane of the ecliptic.

Case C - Radiating surface upright on the equator of the Moon with no solar radiation incident on its surface. The radiator surface is in the plane of the ecliptic.

The coolant temperature range considered in the analyses of these three radiators is at the level encountered when considering the waste-heat radiators for lunar-based power-plant systems presently being considered. The range also covers component coolant loops for these systems, as well as refrigeration system condensing radiators. In the calculations given in this report, the coolant temperature range is from 560° to 1660° R (311 to 922 K). The basic heat-rejection data are presented for a complete lunation for a range of average radiator surface temperatures and the corresponding surface heat-transfer characteristics (ϵ and α_s).

Graphical solutions for equations which convert representative fluid inlet and outlet temperatures to average radiator surface temperatures are also presented. These graphs enhance the usefulness of the information presented herein for approximating radiator areas for various lunar applications.

It must be noted that in this analysis the heat-transfer area is the prime area. An example of prime radiator area would be the two large flat surfaces on a thin rectangular duct in which the heat-rejection fluid is flowing and which have an unobstructed view of the heat sink. In other constructions a multiplication factor including fin effectiveness and interradiation effects, which are dependent on the specific geometry of the radiator, is required in order to obtain actual radiator areas.

METHOD OF ANALYSIS

Heat-Rejection Equations

In the discussion that follows, in order to illustrate the advantage of a vertical radiator over a horizontal radiator, which necessarily has one insulated surface, q will be the heat-rejection flow rate of a 1- by 1-foot (0.305- by 0.305-m) section of radiator regardless of whether it has one or two active sides. The symbol A_R , represents the number of these 1- by 1-foot (0.305- by 0.305-m) radiator sections, and A_R , will represent the actual radiating surface in square feet. In the case of a 1- by 1-foot (0.305- by 0.305-m) section, A_R for a vertical radiator would be 2 square feet (0.186 m²), and for a horizontal radiator 1 square foot (0.093 m²). In both cases only prime radiator area or size is considered.

The rate of heat rejection by the radiator is given by the following equation:

$$Q = wC_p(T_{fi} - T_{fo})$$
 (1)

(All symbols used in this report are defined in appendix A.)

The prime radiating surface area required for a radiator to reject a given amount of heat is determined by the equation (see ref. 1).

$$A_{R}' = wC_{p} \left\{ \frac{1}{h_{R}} \ln \frac{T_{wi}^{4} - T_{s}^{4}}{T_{wo}^{4} - T_{s}^{4}} + \frac{1}{4\sigma\epsilon T_{s}^{3}} \left[\ln \frac{(T_{wi} - T_{s})(T_{wo} + T_{s})}{(T_{wo} - T_{s})(T_{wi} + T_{s})} - 2\left(\arctan \frac{T_{wi}}{T_{s}} - \arctan \frac{T_{wo}}{T_{s}} \right) \right] \right\}$$
(2

If equation (1) is divided by equation (2), the radiator heat flow per unit area is obtained.

$$\frac{Q}{A_{R}^{'}} = \frac{T_{fi} - T_{fo}}{\frac{1}{h_{R}} \ln \frac{T_{wi}^{4} - T_{s}^{4}}{T_{wo}^{4} - T_{s}^{4}} + \frac{1}{4\sigma\epsilon T_{s}^{3}} \left[\ln \frac{(T_{wi} - T_{s})(T_{wo} + T_{s})}{(T_{wo} - T_{s})(T_{wi} + T_{s})} - 2 \left(\arctan \frac{T_{wi}}{T_{s}} - \arctan \frac{T_{wo}}{T_{s}} \right) \right] } .$$
(3)

The fluid temperatures T_{fi} and T_{fo} are normally the initial parameters computed for the radiator. The relations between these temperatures and the radiator wall temperatures T_{wi} and T_{wo} can be obtained by taking a heat balance at any location on the radiator. The heat balance is made between the radiative and convective heat flow for a unit area. The temperature drop through the wall is assumed to be small. Therefore, the outside wall temperature is approximately equal to the inside wall temperature. Equating the convective and the radiative heat flow results in the following expression:

$$h_{R}(T_{f} - T_{w}) = \sigma \epsilon \left(T_{w}^{4} - T_{s}^{4}\right)$$
(4)

Rearranging this equation gives the general expression for the wall temperature

$$T_{\mathbf{W}} = T_{\mathbf{f}} - \frac{\sigma \epsilon}{h_{\mathbf{R}}} \left(T_{\mathbf{W}}^{4} - T_{\mathbf{S}}^{4} \right)$$
 (5)

For the radiator inlet equation (5) becomes

$$T_{wi} = T_{fi} - \frac{\sigma \epsilon}{h_R} \left(T_{wi}^4 - T_s^4 \right)$$
 (6)

For the radiator outlet equation (5) becomes

$$T_{wo} = T_{fo} - \frac{\sigma \epsilon}{h_R} \left(T_{wo}^4 - T_s^4 \right)$$
 (7)

In the equations (4) to (7) the radiator surface emittance ϵ and the convection heat-transfer coefficient h_R can usually be estimated. The effective sink temperature can be calculated as shown in the next section.

Effective Sink Temperature

Lunation, radiator location, and orientation are the variables that affect the sink temperature T_s of the radiator. The radiator sees three different sinks: (1) the surface of the Moon, which is assumed to be black body, (2) space, and (3) the Sun. For this reason equations must be developed expressing the effective sink temperature T_s as a function of the preceding variables. The heat-rejection rate for a 1- by 1-foot (0.305- by

0.305-m) section of radiator with one active side can be expressed in terms of radiator orientation, location, and lunation as

$$q = \epsilon \sigma T_{w,av}^{4} - F_{1} \epsilon \sigma T_{M}^{4} - \left| G_{s} \alpha_{s} \cos \theta_{n} \right|$$
 (8).

The Moon's temperature T_M is given in figure 1. The term G_s is the solar con-

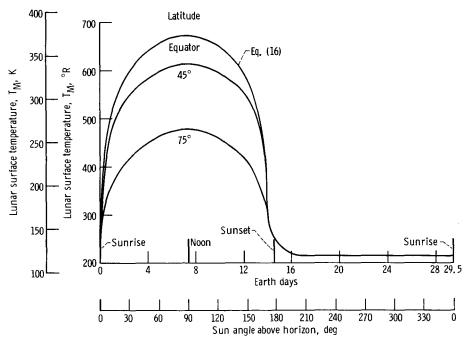
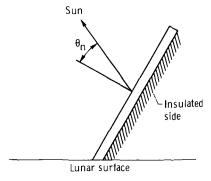


Figure 1. - Lunar surface temperature.

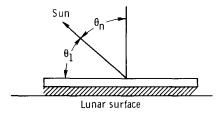
stant, and α_s is the solar absorptance of the radiator surface. The angle of incidence of the Sun is accounted for by θ_n (see fig. 2). Equation (8) as stated above is for one active side and is given in reference 2. The view factor F_1 is for the active side and is given in reference 3. The Moon's albedo is neglected because it is considered a black body (see appendix B). The effect of the presence of the radiator on the Moon's surface temperature is neglected, also.

If the radiator has two active sides, the heat-rejection equation becomes

$$q = 2\epsilon\sigma T_{w,av}^{4} - F_{1}\epsilon\sigma T_{M}^{4} - F_{2}\epsilon\sigma T_{M}^{4} - \left| G_{s}\alpha_{s} \cos \theta_{n} \right|$$
 (9)



(a) General case - one side insulated.



(b) Case A - horizontal radiator.

Figure 2. - Orientation of radiator on surface of Moon.

But $F_1 + F_2 = 1.0$, so equation (9) becomes

$$q = 2\epsilon\sigma T_{w,av}^{4} - \epsilon\sigma T_{M}^{4} - \left|G_{s}\alpha_{s}\cos\theta_{n}\right|$$
 (10)

The radiator average wall temperature $T_{w,av}$ in these equations is the integrated average of the temperature of the radiator surfaces. The method of determining this will be given in a later section.

The heat-rejection rate for a 1- by 1-foot (0.305- by 0.305-m) section of a radiator with one active side can be expressed in terms of average wall and sink temperature by the following equation (The right side of equation (11) happens to be the heat-rejection rate per unit prime area, also):

$$q = \epsilon \sigma \left(T_{w,av}^4 - T_s^4 \right) \tag{11}$$

And, if the radiator has two active sides, the right side of equation (11) must be multiplied by 2.0.

An expression for effective sink temperature for a radiator with one active side, can be found by combining equations (8) and (11).

$$\epsilon \sigma T_{\mathbf{w}, \mathbf{a} \mathbf{v}}^{4} - \mathbf{F}_{1} \epsilon \sigma T_{\mathbf{M}}^{4} - \left| \mathbf{G}_{\mathbf{s}} \alpha_{\mathbf{s}} \cos \theta_{\mathbf{n}} \right| = \epsilon \sigma \left(\mathbf{T}_{\mathbf{w}, \mathbf{a} \mathbf{v}}^{4} - \mathbf{T}_{\mathbf{s}}^{4} \right)$$
 (12)

Solving for the effective sink temperature yields

$$T_{s} = \left[F_{1} T_{M}^{4} + \left| \frac{G_{s}}{\sigma} \left(\frac{\alpha_{s}}{\epsilon} \right) \cos \theta_{n} \right| \right]^{1/4}$$
(13)

For a radiator with two active sides, equation (10) is combined with equation (11). Note that equation (11) must be multiplied by 2.0 because there are two active sides.

$$2\epsilon\sigma T_{\mathbf{w},\mathbf{a}\mathbf{v}}^{4} - \epsilon\sigma T_{\mathbf{M}}^{4} - \left|G_{\mathbf{s}}\alpha_{\mathbf{s}}\cos\theta_{\mathbf{n}}\right| = 2\epsilon\sigma \left(T_{\mathbf{w},\mathbf{a}\mathbf{v}}^{4} - T_{\mathbf{s}}^{4}\right) \tag{14}$$

Solving for T_s results in

$$T_{S} = \left[\frac{T_{M}^{4}}{2} + \left|\frac{G_{S}}{2\sigma} \left(\frac{\alpha_{S}}{\epsilon}\right) \cos \theta_{n}\right|\right]^{1/4}$$
(15)

In the equation for T_s , the surface temperature of the Moon T_M at various locations can be obtained from figure 1. These data were taken from reference 4. The numerical value of the solar constant G_s of 430 Btu/(hr)(ft²) (1.356 kW/m²) used in the calculations herein was taken from references 4 (pp. 7-9) and 5 (p. 236).

The value of the temperature of the Moon at its equator can also be obtained from equation (16) for the interval from sunrise to sunset (ref. 2). In the equation, D_a is Earth days, the argument of the sine is in degrees, and the temperature is in degrees Rankine (to get degrees Kelvin the factor 5/9 must be used).

$$T_{M} = 673 \sin^{1/6} \left(\frac{90 D_{a}}{7} \right)$$
 (16)

Radiator Average Wall Temperature

In a section to follow, an alternate procedure for calculating radiator size is given in terms of $T_{w,av}$. This procedure uses a form of equation (11), which was presented in the development of the sink temperature equations. The inputs to this equation are essentially T_s and $T_{w,av}$. Equations for T_s were developed previously. An expression for $T_{w,av}$ will follow.

It should also be noted that equation (11) is used as the basic equation in the analyses of the performance of the radiator. The performance can be more clearly presented with this equation using $T_{w,a,v}$ and T_{c} rather than equation (3).

with this equation using $T_{w,av}$ and T_s rather than equation (3). When the heat-rejection process is one similar to condensing, the fluid temperature is uniform and nearly constant from inlet to exit conditions. For this case, the arithmetic average of the inlet and exit wall temperatures can be used for $T_{w,av}$. However, in a radiator, the temperature of the fluid decreases from inlet to outlet so that an integrated average should be used for this case.

The expression for the radiator average wall temperature can be developed using a modified form of equation (11). If both sides of equation (11) are multiplied by A_R^{\prime} an expression for the rejected heat flow rate is obtained.

$$Q = \epsilon \sigma A_{R}^{\prime} \left(T_{w,av}^{4} - T_{s}^{4} \right)$$
 (17)

Solving equation (17) for $T_{w,av}$ gives

$$T_{w,av} = \left(\frac{1}{\epsilon \sigma} \frac{Q}{A_{R}^{\prime}} + T_{s}^{4}\right)^{1/4}$$
 (18)

Substituting equation (3) into equation (18) gives the following expression for $T_{w,av}$:

$$T_{w,av} = \left[\frac{T_{fi} - T_{fo}}{e^{\sigma \left\{ \frac{1}{h_R} \ln \frac{T_{wi}^4 - T_s^4}{T_{wo}^4 - T_s^4} + \frac{1}{4\sigma \epsilon T_s^3} \left[\ln \frac{(T_{wi} - T_s)(T_{wo} + T_s)}{(T_{wo} - T_s)(T_{wi} + T_s)} - 2 \left(\arctan \frac{T_{wi}}{T_s} - \arctan \frac{T_{wo}}{T_s} \right) \right] \right\}} + T_s^4} \right]^{1/4}$$
(19)

If equations (6) and (7) are solved for T_{fi} and T_{fo} , respectively, and substituted into equation (19) the following expression is formed:

$$T_{w, av} = \frac{T_{wi} - T_{wo} + \frac{\epsilon \sigma}{h_R} \left(T_{wi}^4 - T_{wo}^4 \right)}{\epsilon \sigma \left\{ \frac{1}{h_R} \ln \frac{T_{wi}^4 - T_s^4}{T_{wo}^4 - T_s^4} + \frac{1}{4\epsilon \sigma T_s^3} \left[\ln \frac{(T_{wi} - T_s)(T_{wo} + T_s)}{(T_{wo} - T_s)(T_{wi} + T_s)} - 2 \left(\arctan \frac{T_{wi}}{T_s} - \arctan \frac{T_{wo}}{T_s} \right) \right] + T_s^4} \right\}$$
(20)

The calculation of $T_{w,av}$ from this equation is not straightforward. For this analysis, equation (20) was plotted with the quantity T_{wi} - T_{wo} as the independent variable and with the quantity T_{wo} as the parameter. The emittance ϵ , film coefficient h_R , and the sink temperature T_s , are held constant for a particular family of curves. A typical set of curves will be presented in the section Lunar Radiator Design Curves.

Specific Design Equations

The basic variable for the cases analyzed in this report is the effective sink temperature T_s . The general expression for this temperature was given previously (eqs. (13) and (15)). The following equations are for the three specific cases considered.

Case A - horizontal position. - For this case the radiator is lying flat on the surface of the Moon. The bottom side is thermally insulated from the surface. Thus, there is only one active side, and equation (13) applies. Since the active side of the radiator does not see the Moon's surface, $F_1 = 0$ and $\cos \theta_n$ can be replaced by $\sin \theta_1$, where θ_1 is the angle of the Sun above the horizon (see fig. 2).

Thus, for this particular case equation (13) reduces to

$$T_{S} = \left[\frac{G_{S}}{\sigma} \left(\frac{\alpha_{S}}{\epsilon}\right) \sin \theta_{1}\right]^{1/4}$$
(21)

Case B - vertical radiator, solar energy incident. - For this case the radiator is normal to the Moon's surface. It is also in a plane normal to the plane of the ecliptic. There are two active sides, so equation (15) gives T_s for this case

$$T_{S} = \left[\frac{T_{M}^{4}}{2} + \left|\frac{G_{S}}{2\sigma}\left(\frac{\alpha_{S}}{\epsilon}\right)\cos\theta_{n}\right|\right]^{1/4}$$
(15)

Case C - vertical radiator, no solar energy incident. - For this case the radiator is normal to the Moon's surface. But is is essentially in the same plane as the ecliptic. Note that these conditions can be met only on the equator. There are two active sides, and the solar contribution is zero. Thus equation (15) becomes

$$T_{S} = \frac{T_{M}}{1.189} \tag{22}$$

As noted previously, T_{M} can be taken from figure 1 or equation (16) may be used.

It should be noted that in all three cases the radiators are at the equator, but appendix B discusses how to handle a radiator problem for a polar location.

Design Procedures

An approach that can be used in sizing a lunar radiator with the analysis presented is to use equation (2). The temperatures T_{wi} and T_{wo} can be determined from known fluid temperatures and combined with appropriate values of T_s , ϵ , α_s , h_R , C_p and fluid flow rate to determine prime radiator area. The effective sink temperatures T_s can be for a specific lunar time or vary over a lunation.

Another approach for calculating radiator sizes is to use the expressions containing average wall temperature developed in the previous section. The procedure for this approach follows.

The value of q is expressed in equation (11) as

$$q = \epsilon \sigma \left(T_{w,av}^4 - T_s^4 \right) \tag{11}$$

The radiator average wall temperature $T_{w,av}$ is found using equation (20), and the effective sink temperature T_s is found using equations (15), (21), or (22). The radiator wall inlet T_{wi} and outlet T_{wo} temperatures (needed in eq. (20)) are determined by equations (6) and (7), respectively, from the known fluid temperatures T_{fi} and T_{fo} . The form of these equations suggests that plotting T_{fi} (or T_{fo}) for assumed values of T_{wi} (or T_{wo}) is the easiest way of determining T_{wi} (and T_{wo}) for given values of h_R , ϵ , and T_s .

After q is determined (eq. (11)), the prime radiator size A_R can be calculated using the following equation (The radiator may have either one or two active sides):

$$A_{R} = \frac{Q}{q} \tag{23}$$

The design procedure for determining radiator areas using these equations is summarized as follows:

- (1) Determine the radiator heat load $\,Q\,$ using equation (1) from given values of fluid temperatures, fluid flow rate, and $\,C_{\rm n}^{}.$
- (2) Determine the effective sink temperature T_S using the appropriate equation (eqs. (15), (21), or (22); curves may be used if they are available).
- (3) Determine T_{wi} and T_{wo} using equations (6) and (7), respectively (curves may be used if they are available).

- (4) Determine $T_{w,av}$ using equation (20) (curves may be used if they are available; be illustrated in the sample problem).
 - (5) Determine q using equation (11) or figure 9 for the specific cases.
 - (6) Determine A_{R} using equation (23).

Radiator prime areas can be found after step 3 in this procedure by using equation (2). In order to base the radiator performance on $T_{w,av}$ the additional steps were included.

RESULTS AND DISCUSSION

Performance

Heat-rejection characteristics. - During a complete lunation, the temperature on the surface of the Moon at its equator varies from about 673° R (374 K) at lunar noon to 213° R (118 K) at lunar midnight. This variation decreases as the latitude increases. In the polar region the lunar temperature is essentially constant. The temperature variation along the equator and at various latitudes for a complete lunation is shown in figure 1.

It can be seen from the expressions for q (eqs. (8) and (10)) that the lunar temperature T_M and the radiator average wall temperature $T_{w,av}$ enter these equations to the fourth power. For this reason $T_{w,av}$ and T_M variations may affect the performance of a radiator appreciably. Orientation of the radiator surface also affects the radiant heat interchange between the radiator and the Moon, space, and Sun. The extent of these effects can be demonstrated by examining each component that makes up equation (8). Representative conditions of Moon and radiator temperatures, orientation, and surface radiative heat-transfer characteristics (table I) are substituted in the expression, and the results are presented in table II.

The two radiator surface temperature conditions shown in table II are representative of low and high values (760° and 1460° R; 422 and 811 K) for lunar systems applications. The radiating surfaces are assumed to have coatings with an emmittance ϵ of 0.90 and a solar absorptance $\alpha_{\rm S}$ of 0.08 and 0.75. The vertical and horizontal radiator positions are considered. In both cases one side of the radiator is insulated.

The first term in equation (8) (table II(a)) is the heat per unit area radiating from the radiator surface. This term varies as the fourth power of the radiator wall temperature. Increasing the temperature from 760° to 1460° R (422.1 and 811.0 K) increases the emission from 514.6 to 7009.1 Btu/(hr)(ft²) (1.6 and 22.1 W/m²).

The second term (table II(b) is the heat input into the radiator (per unit area) from the surface of the Moon. This term is also affected by F_1 . The effect of the Moon on the

TABLE I. - SOLAR ABSORPTANCE AND EMITTANCE OF POSSIBLE RADIATOR COATINGS

						Γ					
	Binder	Surface test	e test	Test	Solar absorptance,	orptance,	Emittance,	Recom	Recommended	Coating	Reference
		temperature	rature	exposure	$\alpha_{\mathbf{S}}$	s	Ψ	edwei	temperature	designation	
		9 ₀	¥	time,	Tnitial	Final		rai	range		
		4	1	sun hr				0 R	Ж		
Magnesium oxide		530	294.1		80.0		06.0		1		5 and 6
	Potassium	550	305	1050	.16	0.18	. 85) To 960	To 533	Z93	L
	silicate	955	530	7200	.17	. 33	. 85) () () () () () () () () () (
	Cr-Co-Ni	To 1060	To 589	1 yr	ო.	e.	.91	To 1060	To 589	A193	8 and 9
	spinel										
	Co 1-percent	To 2260	To 1255	a ₅₀₀₀	b.75	-	88.	To 2260	To 1255	!!!!	10
	Zr										

 $^{\rm a}{\rm Cycle}$ tests, $1960^{\rm o}$ to $2260^{\rm o}$ R (1090 to 1255 K). $^{\rm b}{\rm At}$ 1020° R (567 K). Unpublished data from authors of ref. 9.

TABLE II. - EQUATION 8 TERMS COMPRISING LUNAR RADIATOR

HEAT-REJECTION RATE

[Radiator emittance, 0.90; solar absorptance, 0.08 and 0.75.]

(a) First term, surface emission

т	l temperature, v,av	Surface emi $\epsilon \sigma T_{\mathrm{W}_{j}}^{4}$,
o _R	K	Btu/(hr)(ft ²)	kW/m ²
760 1460	422 811.0	514.6 7009.1	1.623 22.107

(b) Second term, lunar heat input

1	nperature, M	Shape factor, F	Lunar hea F ₁ ¢oʻ	- /
^o R	К		Btu/(hr)(ft ²)	kW/m ²
673	374	0.50 (Vertical) 0 (Horizontal)	155, 48 0	0.4904 0
213	118	0.50 (Vertical) 0 (Horizontal)	1,5876 0	0.005007 0

(c) Third term, solar heat input

Solar included angle,		Solar ab	sorptance, α_s	
$ heta_{ ext{n}}, ext{deg}$	0.08	3	0.75	,
	Sola	r heat inpu	t, $G_s^{\alpha} G_s^{\alpha} \cos \theta_n$	
	Btu/(hr)(ft ²)	$kW/(m^2)$	Btu/(hr)(ft ²)	kW/(m ²)
0 90	34.4 0	0.1082 0	322.5 0	1.015 0

radiator is maximum when the radiator is vertical ($F_1 = 0.50$) and is zero when the radiator is horizontal ($F_1 = 0$). The two lunar temperatures presented are for the extreme conditions (lunar noon, 673° R (374 K); and midnight, 213° R (118 K)) on the Moon's equator. The variation in the Moon's surface temperature at this location is about 460° R (256 K). Thus the second term in equation (8) can vary by a factor of about 100 for a vertical radiator on the equator. For a horizontal radiator this term is zero.

As indicated in figure 1, the least variation for this term for a vertical radiator will occur at or near the pole. The lunar surface temperature variation increases with decreasing latitudes.

The third term in equation (8) (table II(c)) is the heat input to the radiator from the Sun. This term is maximum when $\theta_n = 0^0$ and is zero when $\theta_n = 90^0$.

From table I it can be seen that a coating material with $\alpha_{\rm S}$ of approximately 0.08 may be feasible for a radiator wall temperature of approximately 760° R (422 K), but, for a 1460° R (811 K) material, $\alpha_{\rm S}$ becomes 0.75. However, comparison of the ratio of direct solar radiation to the radiator surface emission (table II) gives 34.4/514.6 = 6.7 percent and 322.5/7009.1 = 4.6 percent for $\alpha_{\rm S}$ = 0.08 and 0.75, respectively. These values demonstrate that direct solar heat is a small percentage of the net radiator heat flux and is not a major problem.

Effective sink-temperature variation. - Variation in effective sink temperature during a complete lunation for one horizontal and two vertical radiators is shown in figure 3.

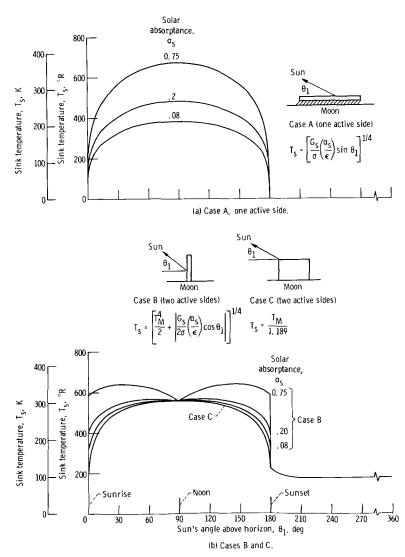


Figure 3. - Effective sink temperature for various radiator positions located on equator. Emittance, 0.90.

For the horizontal case the sink temperature ranges from about 675° to 0° R (375.5 to 0 K) for α_{s} of 0.75. For $\alpha_{s} = 0.08$, the sink temperature ranges from 386° to 0° R (214 to 0 K).

A comparison of these temperatures for the two upright radiators with and without solar radiation incident on the surface shows only a slight difference in magnitude for low α_s 's during the sunrise and sunset hours. At lunar noon, the sink temperature is the same for both upright radiators. For these cases, the sink temperature ranges from about 565° to 180° R (314 to 100.0 K).

The range of fluid temperatures for cooling electronic equipment is from 400° to 570° R (220 to 317 K). Thus a coolant system for such equipment would most likely require a horizontal radiator with a low absorptance coating for operation on the equator over a complete lunation. A vertical radiator will reject heat only during the lunar night-time for these temperatures.

Figure 1 shows that at the higher latitudes the lunar surface temperature decreases. In the polar regions, minimum surface temperatures exist. Deep craters near the poles have permanent shadows on their floors, so in these areas the temperature is constant and near the minimum lunar surface temperatures of approximately 213° R (118 K) (ref. 4). At these locations minimum radiator areas can be achieved because of the low effective sink temperature. A sample problem will be presented in appendix B to show the effect of location on radiator size.

Lunation effects on heat-rejection rates. - The effects of lunation on the heat-rejection capabilities of a radiator are shown in figure 4. The three cases discussed earlier (A, B, and C) are presented for the radiator surface temperatures given in table II (760° and 1460° R or 422 and 811 K). The values of $\alpha_{\rm S}$ used in these calculations were for coatings chosen on the basis of compatability with the radiator surface temperature as indicated in table I. These values are indicated in the figure for each case. The radiators are located on the Moon's equator where the lunation effects are maximum.

Case A - horizontal position: One surface of the radiator is lying flat on the Moon and is thermally insulated from the lunar surface. The radiating surface cannot see the lunar surface, so that \mathbf{F}_1 in equation (8) is zero as shown in table II.

Although the surface temperature of the Moon varies about 460° R (256 K) during a lunation, performance is not greatly affected because the radiator does not see the Moon. For a radiator surface temperature of 760° R (422 K), the heat flux emitted from the radiator is $514.6 \, \text{Btu/(hr)(ft}^2)$ (1.623 kW/m²) (see table II). The solar radiation is zero at lunar sunrise and a maximum of $34.4 \, \text{Btu/(hr)(ft}^2)$ (0.108 kW/m²) at lunar noon for $\alpha_s = 0.08$. Thus the net radiator heat flux is decreased to $480.2 \, \text{Btu/(hr)(ft}^2)$ (1.515 kW/m²). This is a reduction of only about 7 percent. For $\alpha_s = 0.20 \, \text{the reduction increases to 17 percent.}$

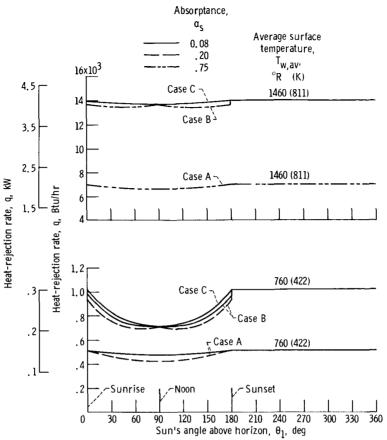


Figure 4. - Effects of Lunation on heat-rejection rate for high and low radiator temperatures. Heat-rejection rate given for 1- by 1-foot (0, 305- by 0, 305-m) section of radiator. Radiators located on equator. Emittance, 0.9. Only prime radiator surfaces considered,

The beneficial effects of increasing radiator wall temperatures for all three cases is clearly shown in figure 4 by the large increase in heat-rejection rate when increasing $T_{w\ av}$ from 760° to 1460° R (422 to 811 K).

Case B - vertical position, solar energy incident: For this case the radiator is normal to the lunar surface, and both sides of the radiator are active. The plane of the radiator is normal to the plane of the ecliptic and is located on the equator. Since both sides of the surface are active, the maximum heat-rejection rate is higher than the horizontal case (case A) by about a factor of two for a comparable radiating temperature.

At the lower radiating temperature (760° R or 422 K) the heat-rejection rate varies by a factor of about 1.44 between lunar midday and night conditions. As the radiating temperature increases (1460° R or 811 K), the midday to night lunar effect decreases to 1.023. This decreasing effect is due to the surface emission of the radiator increasing by the fourth power of temperature while the contribution from the lunar surface and Sun

remains unchanged. The effect on heat-rejection rate for different values of α_s is most apparent during lunar sunrise and sunet.

Case C - vertical position, no direct solar energy incident: For this case the radiator is also normal to the lunar surface and is located on the equator. The plane of the radiator is essentially in the plane of the ecliptic. Thus the direct solar energy is not incident on the surfaces of the radiator. Both sides of the radiator are active. The overall heat-rejection rates for this case are not very different from case B.

Comparison of Cases A, B, and C: Difference in heat-rejection rates for a range of radiator average wall temperatures considered for the three cases are shown in figure 5.

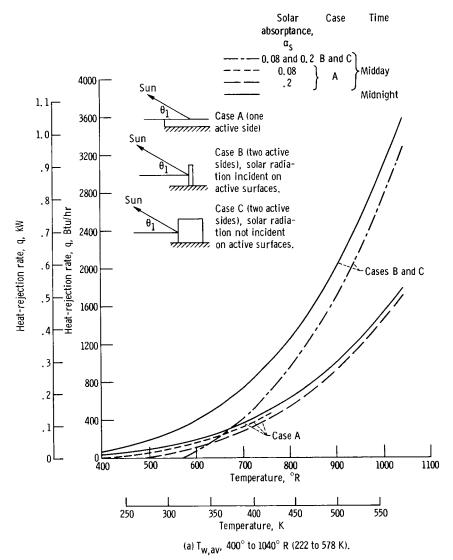
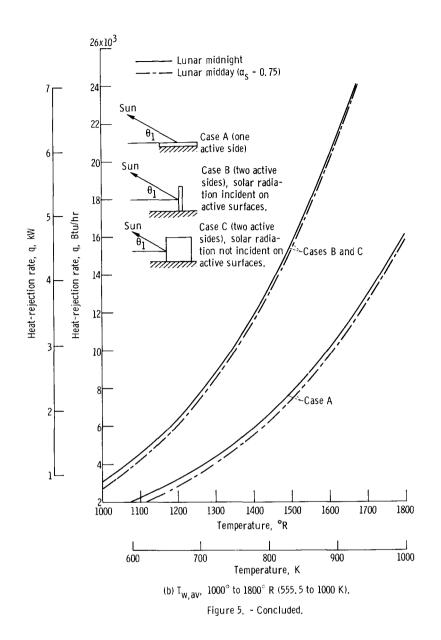


Figure 5. - Heat-rejection rate versus average surface temperature with radiator located on equator. Heat-rejection rate given for 1- by 1-foot (0.305- by 0.305-m) section of radiator. Only prime radiator surfaces considered.



The two specific time periods chosen are lunar high noon and midnight. Figure 5(a) covers $T_{w,av}$ from 400° to 1040° R (222 to 578 K) for $\alpha_s = 0.08$ and 0.20. Figure 5(b) covers $T_{w,av}$ from 1000° to 1800° R (555 to 1000 K) for the same times and $\alpha_s = 0.75$. As noted previously, the results for cases B and C fall on the same curve, and the heat-rejection rate for both cases is about twice the value for case A.

Little difference exists between noon and midnight heat-rejection rates for cases B and C until the design temperature of the radiator falls below 1000^{0} R (555 K). Below this level, the heat-rejection rate for cases B and C during the nighttime period is appreciably higher than noontime.

The magnitude of the effect on heat-rejection rate with appropriate values of α_s between midday and midnight condition can be seen by the results presented in figure 6.

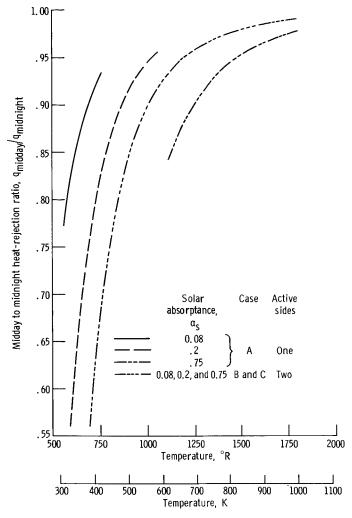


Figure 6. - Midday to midnight heat-rejection ratio versus average surface temperature for radiator located on equator.

In this figure the ratio of midday to midnight heat-rejection rate $q_{midday}/q_{midnight}$ is plotted against the average radiator temperature. The ratio is representative of the variation in performance of a heat-rejection system over a complete lunation at a specific radiator design temperature. All three radiators show less loss of heat-rejection capacity at midday as the radiator temperature increases. For low values of $\alpha_{\rm S}$, the horizontal radiator shows less loss than the vertical. If only coatings with high $\alpha_{\rm S}$ are available, the horizontal radiator suffers greater degradation of capacity than the vertical ones.

Lunar Radiator Design Curves

The procedure used for calculating radiator heat-rejection rates in the performance section indicated the usefulness of design curves for radiator studies. Curves for the solutions of equations (6) or (7), which convert T_f to T_w , and for equation (20), which gives $T_{w,av}$, greatly simplified the computations. In addition, curves like those presented in figure 4 are essentially the solutions of equation (11) as the effective sink temperature varies during one lunation as shown in figure 3.

With three such sets of curves, it is threrfore possible to determine the heat-rejection capacity of a lunar radiator. Figures 7 and 8 present the curves for equations (6) or (7) and (20) for three different values of h_R . Figure 9 presents the solution curves for equation (11) for the three radiator configurations previously considered.

Briefly, the procedure for obtaining heat-rejection rates and radiator size using the design curves is as follows:

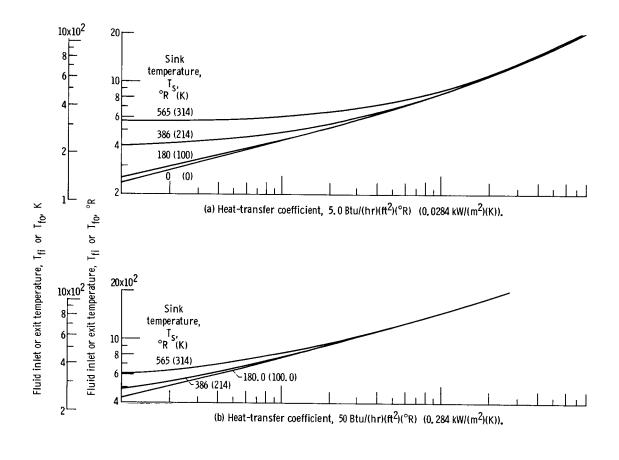
- (1) For a given radiator inlet and outlet fluid temperature T_{fi} and T_{fo} find T_{wi} and T_{wo} using figure 7. Choose the value of h_R most likely to exist in the radiator for the fluid used. The sink temperature T_s can be found from figure 3 for each configuration at any time during the lunar day.
- (2) With T_{wi} and T_{wo} find $T_{w,av}$ using figure 8. Use the same values for h_R and T_s as given previously.
- (3) Knowing $T_{w,av}$, obtain q from figure 9 for the specified orientation, solar absorptance, and the desired lunar time.
- (4) For a given radiator load Q (eq. (1)) and q (determined previously), the prime radiator area $A_{\rm R}$ can be found using equation (23) as shown in the section METHOD OF ANALYSIS:

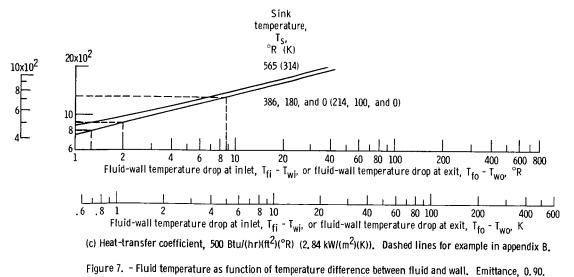
$$A_{R} = \frac{Q}{q}$$

(5) When the actual radiator tube configuration is selected, the heat-transfer coefficient h_R should be checked against the assumed value of h_R for reasonable agreement. A detailed example using these curves will follow in appendix C.

If it becomes advantageous to use different coatings in different temperature regions of a radiator, the above procedure can be used treating the regions as separate radiators in series.

The values of the parameters used in generating these curves are compatible with the systems being considered for lunar application. The radiator average wall temperature ranges from 560° to 1660° R (311 to 922 K) to cover present-day Brayton or Rankine cycle system design. Emittance values of 0.70 and 0.90 were considered during the anal-





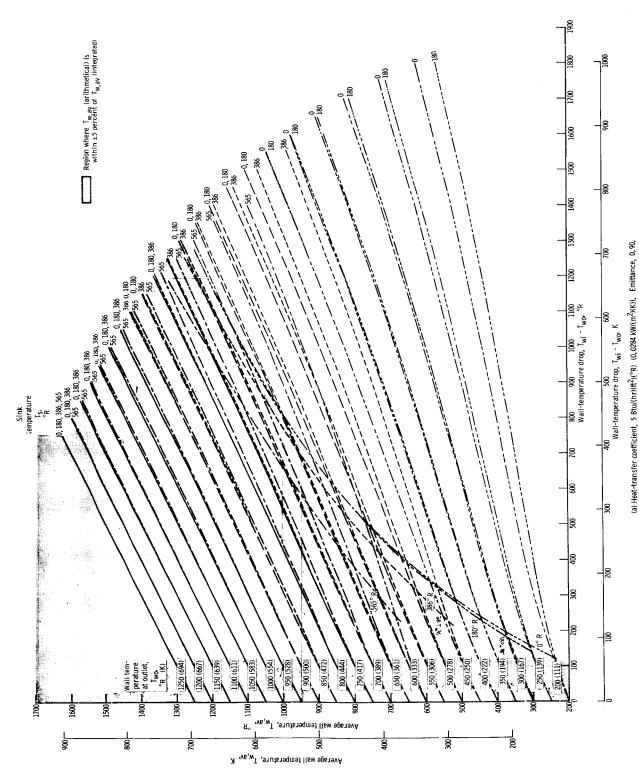
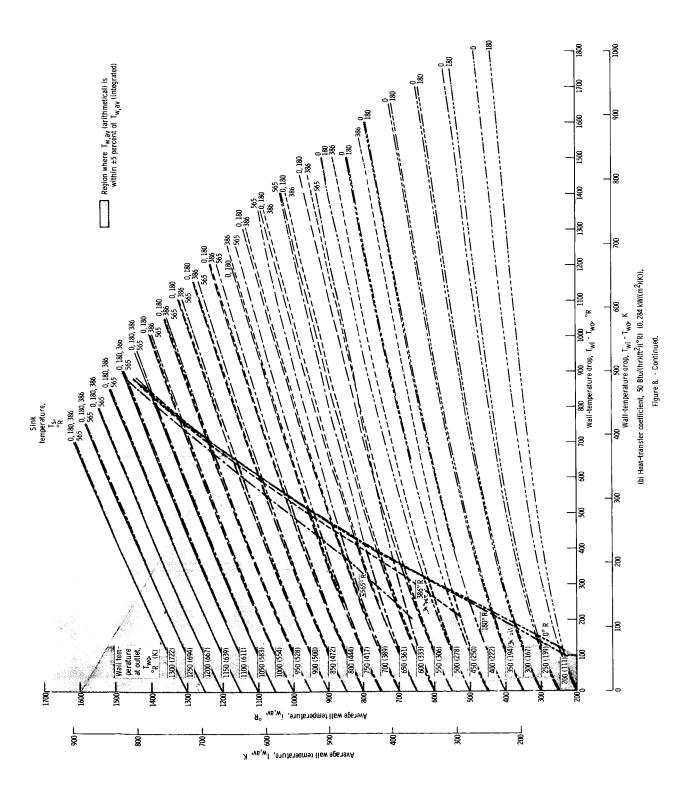
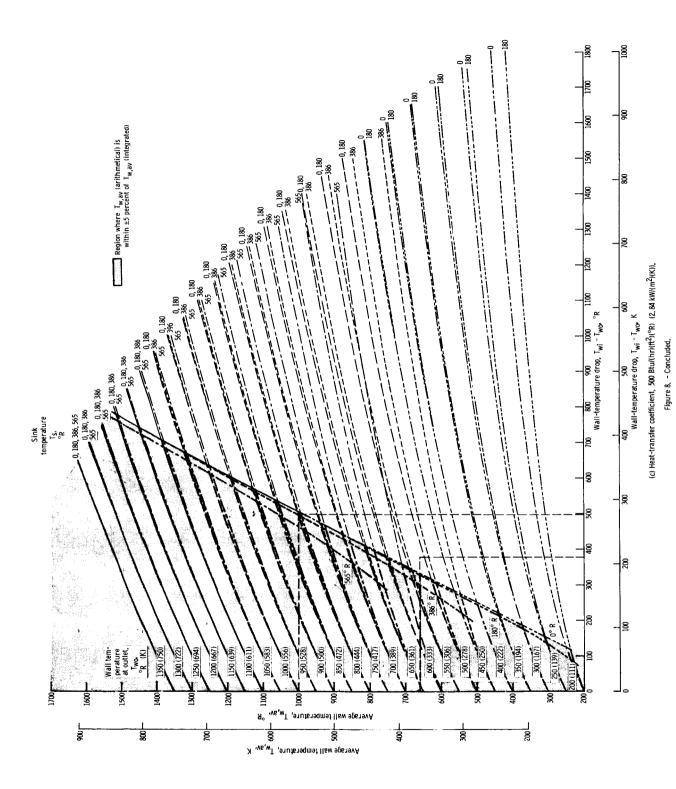


Figure 8. - Average wall temperature as function of wall-temperature drop. (Sink temperature, 0°, 180°, 386°, and 565° R (0, 100, 214, and 565 K.)





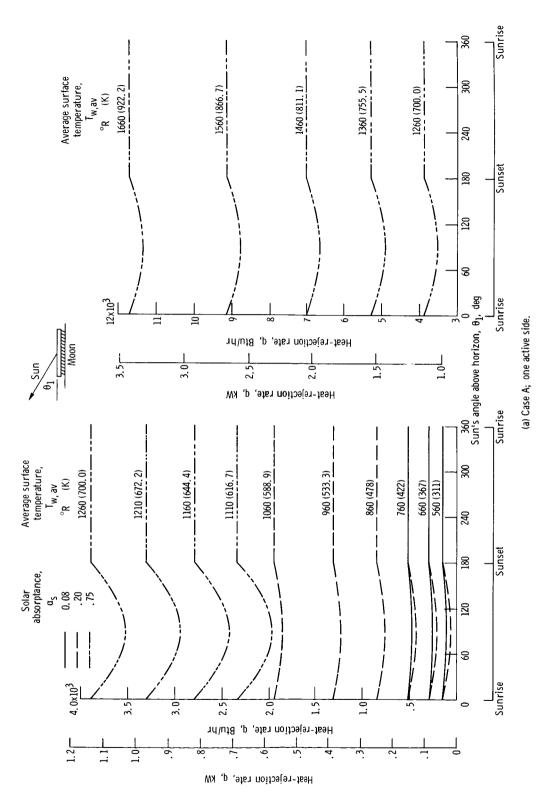
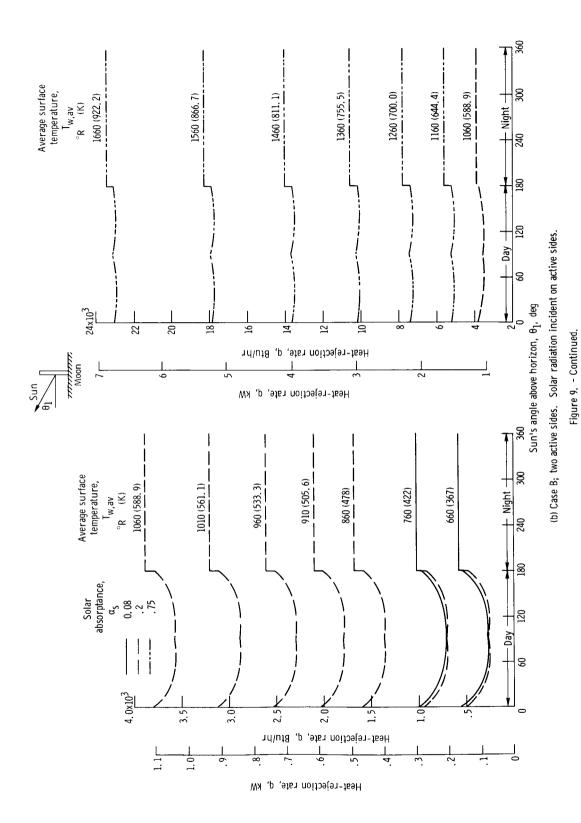
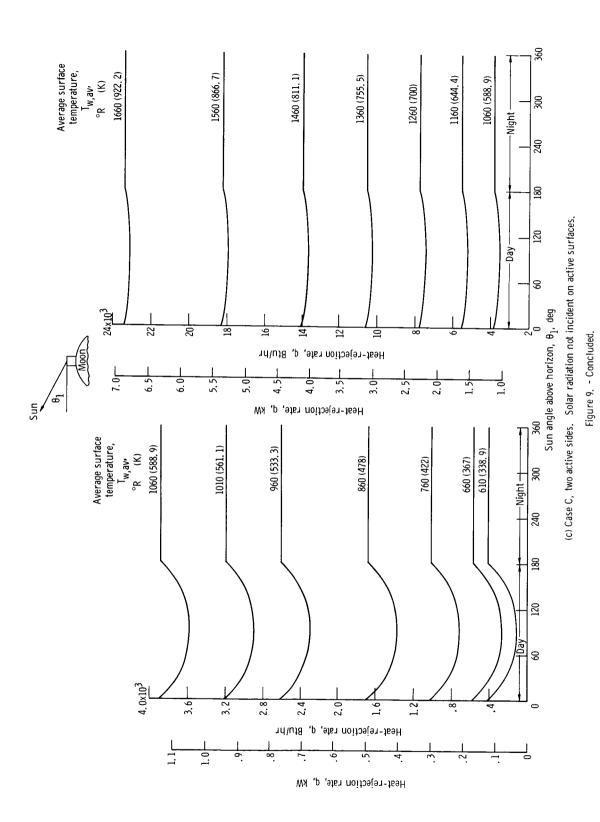


Figure 9, - Radiative heat-rejection rate versus Sun's angle for 1 lunar day with radiator located on equator. Emittance, 0.9; values of q given for 1- by 1-foot (0,305- by 0,305-m) section of radiator. Only prime radiator surfaces considered.





ysis. Since these two values made small changes in the results, only $\epsilon = 0.90$ is shown on these curves.

In the solutions of equations (6) or (7) and (20) (see figs. 7 and 8) three heat-transfer coefficients h_R were selected: 5, 50, and 500 Btu/(hr)(ft²)(0 R) (0.0284, 0.284, and 2.84 kW/(m²)(K)). It was determined that values greater than 500 Btu/(hr)(ft²)(0 R) (2.84 kW/(m²)(K)) had little effect on the results. A value of 500 Btu/(hr)(ft²)(0 R) (2.840 kW/(m²)(K)) is representative of a liquid-metal coolant system. The value of 5Btu/(hr)(ft²)(0 R) (0.0284 kW/(m²)(K)) is representative of a gas system, and 50 Btu/(hr)(ft²)(0 R) (0.0284 kW/(m²)(K)) is for an intermediate condition. The effective sink temperatures used for the curves cover the maximum and minimum values (lunar noon and midnight) for most of the conditions considered for the three radiator cases. For temperatures other than these interpolation is necessary.

The plots of the design curves show certain trends that may be helpful as guide lines in using these curves. The diminishing effect of the effective sink temperature T_s is shown in figure 7 as h_R increases from 5 to 500 Btu/(hr)(ft²)(OR) (0.0284 to 2.840 kW/(m²)(K)).

In general, figure 8 shows at both high radiator outlet temperatures and small $(T_{wi} - T_{wo})$ the effect of T_s on $T_{w,av}$, is negligible. Included in this figure are shaded regions in which the values of $T_{w,av}$ obtained by using an arithmetic average are within 5 percent of the values obtained from equation (20). These regions are presented so that the user can make a quick judgment as to whether an arithmetic average will suffice rather than using equation (20) for points of interest falling away from the parametric values presented. Boundaries for these regions are basically the ordinate and the effective sink temperature. Four specific temperatures considered in the performance studies are shown as the outer boundaries for these regions.

CONCLUDING REMARKS

The equations and graphs in this report permit a rapid estimate of the effective sink temperature and of the size of a waste-heat-rejection radiator for a given lunar thermodynamic system whether for refrigeration, power generation, or component cooling. The analysis herein was limited to polar and equatorial locations and to one horizontal and two vertical attitudes (one in the plane of the ecliptic and one perpendicular to it). The results, however, provide good coverage because they are the extremes of limits.

The general conclusions are

- 1. A vertical radiator is better than a horizontal radiator, at least if its temperature is above 566° R (314 K), because it has two active sides.
 - 2. A polar location is better than an equatorial location.

- 3. High average radiator temperature is desirable from the standpoint of minimizing radiator size.
- 4. Direct solar radiation has little effect, since it appears that a properly selected coating (α_s/ϵ ratio) can match radiator average temperature.

A final choice of radiator geometry, location, and orientation will depend on many additional factors not herein considered: among them, weight of meteoroid protection, radiator structural weight, flexibility of location on lunar surfaces, more accurate accounting of the lunar radiation environment, and operational factors associated with lunar landing, transport, erection and maintenance.

More specific details follow:

5. Direct solar radiation can be made small by proper choice of a α_s/ϵ ratio that is compatable with the radiator average wall temperature, for example:

Radi		Solar absorptance,	Emittance, ϵ	Ratio = M	laximum solar Heat emiss	
o _F	K	$lpha_{\mathbf{s}}$		Percent	Btu/(hr)(ft ²)	kW/m ²
760	422	0.08	0.9	6.7	34.4 514.6	$\frac{0.1082}{1.623}$
1460	811	0.75	0.9	4.6	$\frac{322.5}{7009.1}$	$\frac{1.015}{22.107}$

- 6. If an average radiator wall temperature lower than 566° R (314 K) is required, a horizontal radiator with an insulated back is necessary for daytime operation. The effective sink temperature at high noon on the lunar equator for a vertical radiator is 566° R (314 K). And for a horizontal radiator the effective sink temperature is 386° R (214 K) for an $\alpha_{\rm S}$ = 0.08 and 484° R (268 K) for $\alpha_{\rm S}$ = 0.20.
- 7. In the sample design problem in appendix C, a Brayton cycle powerplant operating in the polar region (sink temperature, $\approx 180^{\circ}$ R or 100 K) is compared with one operating at the equator at lunar midday (sink temperature, $\approx 565^{\circ}$ R or 314 K). The results indicate that a 23-percent reduction in radiator area can be achieved by using the polar location if constant power is required. It should be noted that, in a complete analysis, the effects of changes in efficiencies of other components in the system must be considered.

Lewis Research Center,

National Aeronautics and Space Administration, Cleveland, Ohio, June 22, 1970, 120-27.

APPENDIX A

SYMBOLS

A_{R}	prime radiator size ¹ , ft ² ; m ²	$T_{\mathbf{s}}$	effective sink temperature, ^O R; K
A' _R	prime radiator area, ft ² ; m ²	w	fluid flow rate, lbm/hr; kg/hr
a	albedo	$\alpha_{\mathbf{s}}$	solar absorptance
C _p	specific heat of fluid, Btu/(lb)(OR);	γ	ratio of specific heats
P	$\mathrm{J/(kg)(K)}$	ϵ	emittance of radiator
D_a	Earth days measured from lunar sunrise	$^{\eta}{ m cy}$	Brayton cycle efficiency
F	shape factor or view factor	$^{ heta}$ 1	angle between Sun and horizontal radiating surface
$G_{\mathbf{s}}$	solar constant, 430 Btu/(hr)(ft ²); 1.356 kW/m ²	$\theta_{\mathbf{n}}$	angle between Sun and normal to radiating surface
^h R	convective heat-transfer coefficient based on outside radiating prime area, Btu/(hr)(ft ²)(OR); kW/(m ²)(K)	σ	Stefan-Boltzmann Constant, $0.17132\times10^{-8} \text{ Btu/(hr)(ft}^2)(^{0}R^4);$ $5.6671\times10^{-11} \text{ kW/(m}^2)(K^4)$
	17.44 / / 1111 / / 17/		
т	, , , , , ,	Subsci	ripts:
I _n	solar energy rate reflected per Moon unit area, Btu/(hr)(ft ²);	Subscr	ripts: albedo
I _n	solar energy rate reflected per		
I _n	solar energy rate reflected per Moon unit area, Btu/(hr)(ft ²);	a	albedo
	solar energy rate reflected per Moon unit area, Btu/(hr)(ft ²); kW/(m ²)	a av	albedo integrated average
	solar energy rate reflected per Moon unit area, Btu/(hr)(ft ²); kW/(m ²) total heat-rejection flow rate, Btu/hr; kW heat-rejection flow rate for a	a av f	albedo integrated average fluid
Q	<pre>solar energy rate reflected per Moon unit area, Btu/(hr)(ft²); kW/(m²) total heat-rejection flow rate, Btu/hr; kW</pre>	a av f i	albedo integrated average fluid inlet
Q q	solar energy rate reflected per Moon unit area, Btu/(hr)(ft ²); kW/(m ²) total heat-rejection flow rate, Btu/hr; kW heat-rejection flow rate for a 1- by 1-ft (0.305- by 0.305-m) section of radiator ¹	a av f i M	albedo integrated average fluid inlet Moon
Q	solar energy rate reflected per Moon unit area, Btu/(hr)(ft ²); kW/(m ²) total heat-rejection flow rate, Btu/hr; kW heat-rejection flow rate for a 1- by 1-ft (0.305- by 0.305-m)	a av f i M o	albedo integrated average fluid inlet Moon outlet
Q q	solar energy rate reflected per Moon unit area, Btu/(hr)(ft ²); kW/(m ²) total heat-rejection flow rate, Btu/hr; kW heat-rejection flow rate for a 1- by 1-ft (0.305- by 0.305-m) section of radiator ¹ solar albedo energy rate being absorbed by a 1- by 1-ft (0.305- by 0.305-m) section	a av f i M o R	albedo integrated average fluid inlet Moon outlet radiator
Q q ^q a	solar energy rate reflected per Moon unit area, Btu/(hr)(ft ²); kW/(m ²) total heat-rejection flow rate, Btu/hr; kW heat-rejection flow rate for a 1- by 1-ft (0.305- by 0.305-m) section of radiator ¹ solar albedo energy rate being absorbed by a 1- by 1-ft (0.305- by 0.305-m) section of vertical radiator, Btu/hr; kW	a av f i M o R rej	albedo integrated average fluid inlet Moon outlet radiator rejection wall active side of a radiator with one
Q q	solar energy rate reflected per Moon unit area, Btu/(hr)(ft ²); kW/(m ²) total heat-rejection flow rate, Btu/hr; kW heat-rejection flow rate for a 1- by 1-ft (0.305- by 0.305-m) section of radiator ¹ solar albedo energy rate being absorbed by a 1- by 1-ft (0.305- by 0.305-m) section	a av f i M o R rej w	albedo integrated average fluid inlet Moon outlet radiator rejection wall

¹Radiator may have one active side or two active sides; also only prime radiator area or size is considered.

second side of a radiator with two active sides

Superscript:

- * Brayton cycle temperatures (see fig. 11)
- ' heat exchanger fluid temperatures, waste heat loop (see fig. 11)

APPENDIX B

THE MOON'S ALBEDO

For case A, the radiator is lying flat on the Moon, and the bottom is thermally insulated from the Moon's surface. Since the surface of the radiator does not see the surface of the Moon, there is no albedo heat flux incident on the radiator.

For the analysis of both cases B and C, the curvature of the lunar surface to be a plane. Also, it is commonly assumed that the planets and the Moon have diffusely reflecting surfaces which obey Lambert's Cosine Law (see ref. 11). Therefore, the shape or view factor (ref. 3) for a vertical radiator with two active sides will be 1.0 for both cases B and C.

The solar energy rate reflected per Moon unit area is given by the equation

$$I_r = G_S a \sin \theta_1 \tag{B1}$$

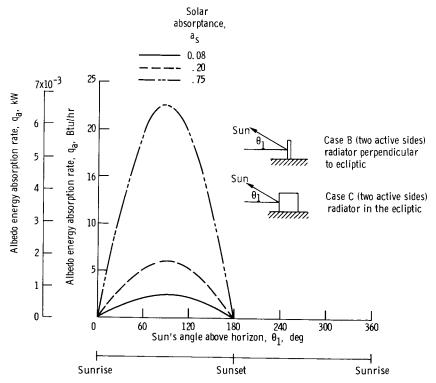


Figure 10. - Albedo energy absorption rate versus Sun's angle for 1 lunar day with radiator located at equator for cases B and C. Ordinate for a 1- by 1-foot (0, 305- by 0, 305-m) section of radiator.

The solar energy rate being absorbed by a 1- by 1-foot (0.305- by 0.305-m) section of vertical radiator is

$$q_a = F_a \alpha_s I_r = F_a \alpha_s G_s \text{ a sin } \theta_1$$
 (B2)

where F_a is 1.0 as stated albedo, G_s is 430 Btu/(hr)(ft²) (1.356 kW/m²), and the albedo a from reference 11 is 0.07. Figure 10 is a plot of the albedo energy absorption rate for the lunar daylight hours for a 1- by 1-foot (0.305- by 0.305-m) section of vertical radiator in which the several values of solar absorptance α_s used in the report are parameters. These curves can be used to make corrections to the data given in the report. However, from equation (B2), it is apparent that, even for the maximum albedo values, the energy rate incident upon the radiator is less than one tenth the value of the maximum direct energy rate from the Sun, and in most cases can be neglected.

APPENDIX C

SAMPLE PROBLEM USING DESIGN CURVES

To demonstrate the use of the curves two sample calculations will be performed. These calculations show how the radiator area is affected by the differences in sink temperature at the equator and at the poles of the Moon.

Two Brayton cycle power conversion systems are to be analyzed, both having the following conditions:

Alternator efficiency
Electrical power output, kW
Turbine inlet temperature, ^O R (K)
Turbine efficiency
Turbine-to-compressor pressure loss ratio
Recuperator effectiveness
Compressor efficiency

The combined temperature-entropy and flow diagram is shown in figure 11.

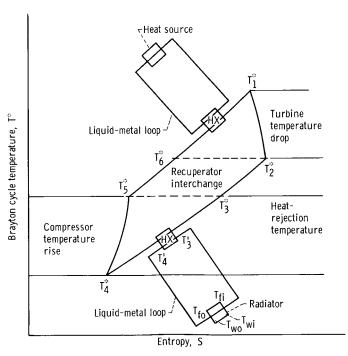


Figure 11. -Working temperatures for Brayton cycle power conversion system.

One system is located in the polar region with an effective sink temperature of approximately 180° R (100 K), which is also approximately the nighttime sink temperature at the equator. The other is located at the equator with a maximum sink temperature of 565° R (314 K) (lunar noon). In each case it is assumed that full-power operation is required throughout the lunar cycle. The systems will have a vertical radiator with two active sides. This configuration gives the maximum heat-rejection per unit area. An optimization procedure given in reference 1 will be used to obtain minimum radiator area for each case. The procedure involves plotting radiator areas against cycle temperature ratios for different turbine temperature ratios. Figures 7 to 9 will be used to obtain the radiator areas for these computations.

Because it is desirable to exclude the complication, the drop through the walls of the heat-rejection rate were also calculated. An example of this information is given Therefore,

$$T_3^{\dagger} = T_3^* = T_{fi}$$

and,

$$T_4' = T_4^* = T_{fo}$$

With the use of the equations given in reference 1, and with a γ (ratio of specific heats) of 1.67, the state points of the working fluid around the Brayton cycle were calculated for the various values of the parameters, cycle temperature ratio $(T_4/T_1)^*$, and turbine temperature ratio $(T_2/T_1)^*$. The thermodynamic cycle efficiency η_{cy} and the system heatrejection rate were also calculated. An example of this information is given in table III for the turbine temperature ratio $(T_9/T_1)^*=0.75$.

The following table is abstracted from table III and shows, for the two values of $(T_4/T_1)^*$, the various derived values and the method of geeting them.

1	2	2	3		4		6		6		7		8			9
Cycle					_			Source o	r deriva	ition						
temper- ature ratio, (T_4/T_1)	From table I fig.	7(c)	From table II fig. '	II and 7(c)	Colur minus in tab	s T ₃	Colui minus in tab	т ₄	Colui min colur	ius	From fi using co 5 and 6	lumns	From co		rate divid	ejection Q _{rej} ded by mn 8
	T _{fi} -	Twi	T _{fo} - T _{wo}		Twi		T	vo	T _{wi} -	Two	Т _w ,	av	g	Į.	A	'R
	OR	К	o _R	К	o _R	К	^o R K		^o R	К	^o R K		Btu/hr kW		ft ²	m ²
0.25	2.0	1.1	<1	<0.51	893.0	496.1	515.0	286.0	378.0	210.1	660.0	366.7	582	0.171	5962	552.5
. 39	8.8	4.9	1.25	.69	1300.2	722.3	801.8	445.3	498.4	277.0	1000.0	555.6	3080	. 902	8858	824.0

TABLE III. - BRAYTON CYCLE RADIATOR AREA CALCULATION

[Brayton system power output, 500 kW_e; turbine inlet temperature, 2060° R (1144 K); Fink temperature, 180° R (100 K); turbine temperature ratio, 0.75; recuperator effectiveness, 0.90; radiator heat-transfer coefficient, 500 Btu/(hr)(ft²)($^{\circ}$ R) (2.839 kW/(m²)(K)).]

(a) U.S. Customary Units

																			_		\neg
Prime	radiator	size	¥		11	2962	5 726	5 360	5 142	4 859	4 720	4 630	4 616	4 795	4 886	5 114	5 455	6 047	066 9	8 8 8 8 8 8 8 8 8 8	12 526
Heat-	rejection	rate,	ö	Btu/hr		582	099	770	880	1026	1170	1330	1500	1640	1850	2065	2310	2560	2840	3080	3410
Average	wall	temper-	ature.	T	w, av. ^o R	099	685	710	735	160	785	810	835	855	880	902	930	922	980	1000	1025
Inlet-exit	wall tem-	perature	drop.	T T	wı wı	378.0	385.8	395.5	404.1	412.7	421.4	429.0	438.6	447.2	455.6	464.0	472.4	481.8	490.1	498.4	506.6
Exit	wall	temper-	ature.	Τ	o R	515.0	536.0	556.0	577.0	597.0	618.0	639.0	659.0	680.0	700.0	721.0	742.0	761.0	781.9	801.8	822.6
Inlet	wall	temper-	ature.	Ţ	o ^R	893.0	921.8	951.5	981.1	1009.7	1039.4	1068.0	1097.6	1127.2	1155.6	1185.0	1214.4	1242.8	1272.0	1300.2	1329.2
Fluid-wall	temper-	ature drop	at exit.	T T.	IO WO.	0											-	1.00	1.10	1, 25	1.40
Fluid-wall	temperature	drop at	inlet.	T T	n wi	2.0	2.2	2.5	2.9	3.3	3.6	4.0	4.4	4.8	5.4	6.0	6.6	7.2	8.0	8.8	9.8
Heat-	rejection	rate,	3	rej' Btu/hr		3.47×10 ⁶	3.779	4.127	4.525	4.985	5.522	6.158	6.924	7,862	9.040	10, 560	12.600	15.480	19.852	27.283	42.712
Cycle	effi-	ciency,	2	, cy		0.353	. 334	.315	. 295	. 276	. 256	. 235	. 215	. 194	.173	. 152	. 131	. 109	. 087	. 065	. 043
				$^{\mathrm{T}}_{6}$	1	1473	1476	1479	1483	1486	1489	1493	1496	1499	1502	1506	1509	1512	1516	1519	1522
tures,				T		823	856	888	921	954	987	1020	1053	1086	1119	1152	1185	1217	1250	1283	1316
mpera				T ₄		515	536	556	577	597	618	639	629	680	700	721	742	762	783		824
Brayton cycle temperatures,	$^{\rm o}$	(a)		T ₃)	895	924	954	984	1013	1043	1072	1102	1132	1161	1191	1221	1250	1280	1309	1339
3raytor				$^{\mathrm{T}_2}$	ı	1545						_							_		-
				Т,		2060															-
Cycle	temper-	ature	ratio	T,/T,	1	0,25	. 26	. 27	. 28	. 29	.30	.31	. 32	. 33	. 34	. 35	. 36	. 37	38	. 39	. 40

^aSee fig. 11 for identification of Brayton cycle temperatures.

TABLE III. - Concluded. BRAYTON CYCLE RADIATOR AREA CALCULATIONS

[Brayton system power output, 500 kW_e; turbine inlet temperature, 2060^o R (1144 K); sink temperature, 180^o R (100 K); turbine temperature ratio, 0.75; recuperator effectiveness, 0.90; radiator heat-transfer coefficient, 500 Bu/(hr)(ft²)(PR) (2.839 kW/(m²)(K)).]

			_															
	Prime size A _D ,	H ₂	552.5	533.0	497.2	477.8	455.0	438.3	430.0	429.8	445.0	454.0	475.2	506.5	562.0	650.0	824.0	1165.0
	Heat- rejection rate,	q, kW	0.171	. 193	. 226	. 258	300	. 343	. 390	. 439	.481	. 542	.605	.677	. 750	. 832	. 902	666.
	Average wall temper-		366.7	380.6	394, 4	408.3	422.2	436.1	450.0	463.9	475.0	488.9	502.8	516.7	530.6	544.4	555.6	569.4
	Inlet-exit wall tem-	$\begin{array}{c} \text{drop,} \\ T_{wi} - T_{wo}, \\ K \end{array}$	210.1	214.0	219.5	224.0	228.9	234.4	238.3	243.7	248.2	252.9	257.3	262.9	268.0	272.2	277.0	281.2
	Exit wall temper-	ature, T _{wo} , K	286.0	298.0	309.0	321.0	332.0	343.0	355.0	366.0	378.0	389.0	401.0	412.0	422.4	434.4	445.3	457.2
	Inlet wall temper-	ature, T _{wi} , K	496.1	512.0	528.5	545.0	560.9	577.4	593.3	609.7	626.2	641.9	658.3	674.9	690.4	706.6	722.3	738.4
	Fluid-wall temper-ature drop	at exit, $T_{fo} - T_{wo},$ K	0				·						•	-	. 56	.61	69.	.78
(b) SI Units	Fluid-wall temperature drop at	inlet, T _{fi} - T _{wi} , K	1.1	1.3	1.4	1.6	1.8	2.0	2.2	4.2	2.7	3.0	3.3	3.7	4.0	4.4	4.9	5.4
	Heat- rejection rate,	Qrej, kW	1 017	1 107	1 209	1 326	1 461	1 618	1 804	2 029	2 304	2 649	3 094	3 692	4 536	5 817	7 994	12 515
	Cycle effi- ciency,	η_{cy}	0.353	. 334	.315	. 295	. 276	. 256	. 235	. 215	. 194	.173	.152	. 131	. 109	. 087	.065	.043
		r ₉	818	820	822	824	825	827	829	831	833	834	837	838	840	842	844	845
	atures,	T	457	476	493	512	530	548	267	585	603	622	640	658	919	694	713	731
	temper:	T ₄	286	298	309	321	332	343	355	366	378	389	401	412	423	435	446	458
	Brayton cycle temperatures, K (a)	T 3	497	513	230	547	563	579	296	612	629	645	662	678	694	711	727	744
	3rayton	T	828		_	_												-
		$^{\mathrm{T}_{1}}$	1144															
	Cycle temper- ature	ratio, ${ m T}_4/{ m T}_1$	0, 25	. 26	. 27	. 28	. 29	. 30	. 31	. 32	. 33	. 34	.35	. 36	. 37	.38	. 39	. 40

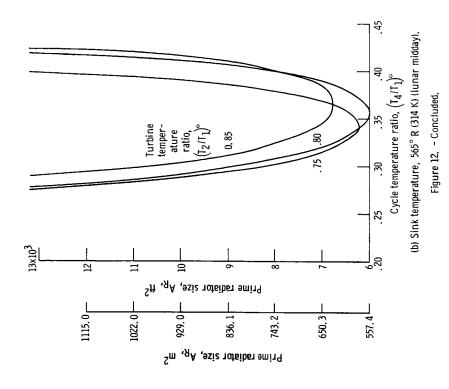
^aSe⁽¹⁾ fig. 11 for identification of Brayton cycle temperature.

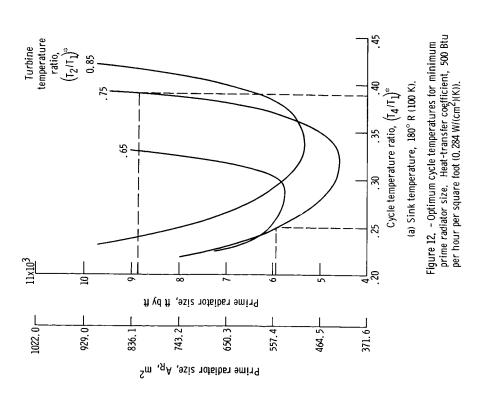
The temperature drop between the liquid metal heat-rejection fluid and the radiator surface, for an assumed $\rm h_R$ of 500.0 Btu/(hr)(ft²)(^OR) (2.840 kW/(m²)(K)), is given in columns 2 and 3 for the inlet and outlet, respectively. As shown in column 2, the data are obtained from figure 7(c) (see the dashed lines); the argument or independent variable is $\rm T_3$ or $\rm T_{fi}$, which is the ordinate, and the temperature drop, $\rm T_{fi}$ - $\rm T_{wi}$, is given by the abscissa. In a similar manner $\rm T_{fo}$ - $\rm T_{wo}$ of column 3 is obtained in which $\rm T_4$ or $\rm T_{fo}$ is used as the argument. All values less than 1 OR (0.156 K) are neglected.

As shown in columns 4 and 5, T_{wi} and T_{wo} are obtained by subtracting T_3 from column 2 and T_4 from column 3, respectively. Column 6, T_{wi} - T_{wo} , is simply the difference between columns 4 and 5.

The values for $T_{w,av}$ in column 7 are obtained from the ordinates of figure 8(c) (see the dashed lines in the figure) in which the parameters are $T_s = 180^{\circ}$ R (100 K) and the wall outlet temperature T_{wo} (column 5). The argument (abscissa) is T_{wi} - T_{wo} , and the two values 378° R (210.1 K) and 498.4° R (277 K) are indicated by the dashed lines on figure 8(c) which lead to the two values of $T_{w,av}$ of 660° R (366.7 K) and 1000° R (555.6 K), respectively. For these values of $T_{w,av}$ (660° and 1000° R or 366.7 and 555.6 K) shown in column 8 the heat-rejection rates for a 1- by 1-foot (0.305- by 0.305-m) section of radiator per unit area q are 582.0 and 3080.0 Btu/hr (0.171 and 0.902 kW), respectively (from fig. 5(a)). Dividing the system heat-rejection rate Q_{rej} (table III) by q, the heat-rejection rate for a 1- by 1-foot (0.305- by 0.305-m) section, gives the prime radiator size A_R . These two points are illustrated in figure 12(a). In a similar manner the prime radiator area is calculated for other values of $(T_4/T_1)^*$ to generate the curve $(T_2/T_1)^* = 0.75$ in figure 12(a). The entire procedure is repeated to produce curves for $(T_2/T_1)^* = 0.65$ and 0.85. The results for a sink temperature of 565° R (314 K) are given in figure 12(b).

The curves (figs. 12(a) and (b)) show that the optimum prime radiator for a sink temperature of 180° R (100 K) is about 4600 square feet (410 m²). For a sink temperature of 565° R (314 K), the prime radiator area is about 6000 square feet (558 m²). These results are indicative of what can be expected at different lunar locations. However, many other system problems would have to be considered before a powerplant is positioned in the polar regions solely for the reduction of radiator size.





REFERENCES

- 1. Glassman, Arthur J.; and Stewart, Warner L.: A Look at the Thermodynamic Characteristics of Brayton Cycles for Space Power. Paper 63-218 AIAA, June 1963.
- 2. Mackay, Donald B.: Design of Space Powerplants. Prentice-Hall, Inc., 1963.
- 3. Hamilton, D. C.; and Morgan, W. R.: Radiant-Interchange Configuration Factors. NACA TN 2836, 1952.
- 4. Niedz, F. J.: Survey of the Physical and Environmental Parameters of the Moon. Rep. 4810, General Electric Co. (NASA CR-55064), Feb. 1963.
- 5. Stevenson, Jack A.; and Grafton, John C.: Lunar Temperature Environment.

 Proceedings of the 1969 National Meeting of the Institute of Environmental Sicences,
 Los Angeles, Apr. 6-8, 1960.
- 6. McAdams, William H.: Heat Transmission. Third Ed., McGraw-Hill Book Co., Inc., 1954, p. 63.
- 7. Grammer, J. R.; and Cunnington, G. R.: Emissivity Coatings for Low-Temperature Space Radiators. Rep. L-23-68-1, Lockheed Missles and Space Co., June 1968.
- 8. Crosby, J. R.: The development and Qualification of Thermal Control Coatings for SNAP Systems. Rep. NAA-SR-9908, Atomics International, Sept. 30, 1965.
- 9. Diedrich, James H.; and Curtis, Henry B.: Experimental Investigation of Total Emittance and Solar Absorptance of Several Coatings Between 300° and 575° K. NASA TN D-3381, 1966.
- 10. Cleary, R. E.; Emanuelson, R.; Luoma, W.; and Ammann, C.: Properties of High Emittance Materials. NASA CR-1278, 1969.
- 11. Ballinger, J.C.; Elizalde, J. C.; Garcia-Varela, R. M.; and Christensen, E. H.: Environmental Control Study of Space Vehicles. Part II: Thermal Environments of Space. Rep. ERR-AN-016, General Dynamics/Convair Astronautics, Nov. 1, 1960.